

Quantum, Classical and Symmetry Aspects of Max Born Reciprocity

L.M. Tomilchik

B.I. Stepanov Institute of Physics,
National Academy of Sciences of Belarus

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Topics

- ① Reciprocal Symmetry and Maximum Tension Principle
- ② Maximum Force and Newton gravity
- ③ Extended Phase Space (QTPH) as a Basic Manifold
- ④ Complex Lorentz group with Real Metric as Group of Reciprocal Symmetry
- ⑤ One-Particle Quasi-Newtonian Reciprocal-Invariant Hamiltonian dynamics
- ⑥ Canonic Quantization: Dirac Oscillator as Model of Fermion with a Plank mass

Born Reciprocity version

- ① Reciprocity Transformations (RT):

$$\frac{x_\mu}{q_e} \rightarrow \frac{p_\mu}{p_e}, \quad \frac{p_\mu}{p_e} \rightarrow -\frac{x_\mu}{q_e}.$$

- ② Lorentz and Reciprocal Invariant quadratic form:

$$S_B^2 = \frac{1}{q_e^2} x^\mu x_\mu + \frac{1}{p_e^2} p^\mu p_\mu, \quad \eta_{\mu\nu} = \text{diag}\{1, -1, -1, -1\}$$

- ③ x_μ, p_μ are quantum-mechanical canonic operators:

$$[x_\mu, p_\nu] = i\hbar\eta_{\mu\nu}$$

- ④ Phenomenological parameter according definition
 $p_e = M\mathbf{c}$ so that

$$q_e = \hbar/M\mathbf{c} = \Lambda_c \quad (\text{Compton length}).$$

The mass M is a free model parameter.

- ⑤ Two dimensional constants q_e [length], p_e [momentum]
connected by correlation:

$$q_e p_e = \hbar \quad (\text{Plank constant}).$$

Relativistic Oscillator Equation (ROE)

Selfreciprocal-Invariant Quantum mechanical Equation:

$$\left(-\frac{\partial^2}{\partial \xi^\mu \partial \xi_\mu} + \xi^\mu \xi_\mu \right) \Psi(\xi) = \lambda_B \Psi(\xi),$$

where $\xi^\mu = x^\mu / \Lambda_c$.

The (ROE) symmetry $\approx U(\mathbf{3}, \mathbf{1})$. There are solution corresponding linear discrete mass spectrum.

Maximum Tension Principle (MTP)

The space-time and momentum variables are to be just the same dimensions of a quantity. It is necessary to have the universal constant with dimension: momentum/length, or equivalently: mass/time, energy/time, momentum/time. In reality: MTP [Gibbons, 2002]

Gibbons Limit:

$$\text{Maximum force} \quad F_{max} = \left(\frac{dp}{dt} \right)_{max} = \frac{c^4}{4G_N} = F_G,$$

$$\text{Maximum power} \quad P_{max} = \left(\frac{dE}{dt} \right)_{max} = \frac{c^5}{4G_N} = P_G.$$

[momentum]/[length] – Universal constant

$$\mathfrak{x}_0 = \frac{p_e}{q_e} = \frac{c^3}{4G_N} \quad [\text{LMT, 1974, 2003}]$$

It is evident: $\mathfrak{x}_0 = c^{-1} F_{max} = c^{-2} P_{max}$. All these values are *essentially classic* (does not contain Plank constant \hbar).

Born Reciprocity and Gibbons Limit

Let the parameter \mathbf{a} to be some fixed (nonzero!) value of action S . We replace initial Born's relation

$$q_e p_e = \hbar \quad \text{by} \quad \pi q_e p_e = \mathbf{a}$$

without beforehand action discreteness supposition. We demand the nonzero area $S_{fix} = \pi q_e p_e$ of the phase plane only. The second suggestion is:

$$\mathfrak{a}_0 = \frac{p_e}{q_e} = \frac{c^3}{4G_N} \quad (\text{universal constant}).$$

Now the parameters q_e , p_e are defined separately

$$q_e = \sqrt{\frac{a}{\pi \mathfrak{a}}}, \quad p_e = \sqrt{\frac{a \mathfrak{a}}{\pi}}$$

Evidently: the phenomenological constants $E_e = p_e c$ [Energy] and $t_e = c^{-1} q_e$ [time] satisfy the conditions

$$E_e t_e = \pi a, \quad \frac{E_e}{t_e} = c^2 \mathfrak{a} = \frac{c^5}{4G_N}$$

Under Born's parametrization

$$p_e = Mc, \quad q_e = \frac{2MG_N}{c^2} = r_g(M) - \text{gravitational radius.}$$

General mass-action correlation

$$M^2 = \frac{ac}{4\pi G_N}.$$

It is valid as in classical as in quantum case. Under supposition when $a_{min} = \hbar$ we receive

$$M^2 = \frac{\hbar c}{4\pi G_N} = \frac{1}{2} M_{Pl}^2, \quad M_{Pl} = \sqrt{\frac{\hbar c}{G_N}} - \text{Plank mass.}$$

Maximum Force and Modified Newton Gravity

$$\begin{aligned} \text{(NG): } |\underline{f}| &= G_N \frac{mM}{r^2} \equiv \frac{c^4}{4G_N} \frac{2mG_N}{c^2} \frac{2MG_N}{c^2} \frac{1}{r^2} \\ &= F_G \frac{r_g R_g}{r^2} \equiv F_G \frac{r_0^2}{r^2} \quad (r \geq 0, \quad r_0 = \sqrt{r_g R_g}) \end{aligned}$$

$$\text{(ModNG): } |f_N^{mod}| = F_G \frac{r_0^2}{r^2} \quad (r \geq r_0), \quad |f_N^{mod}|_{r=r_0}^{max} = F_G$$

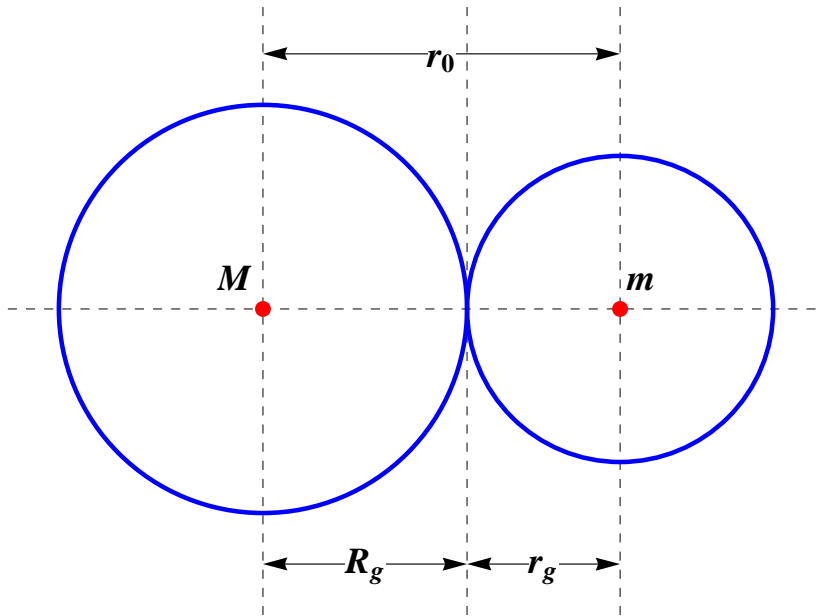
The corresponding gravitational potential Energy

$$U_{gr}^{mod} = -F_G \frac{r_0^2}{r}$$

possess a minimum under $r = r_0$

$$U_{gr}^{mod}(min) = -F_G r_0 = -\frac{c^4}{4G_N} \frac{2G_N}{c^2} \sqrt{mM} = -\frac{c^2}{2} \sqrt{mM}$$

We have dealing (instead of point-like masses) with a pair spacely extended simple massive gravitating objects like the elastic balls or drops with a high surface tension.



$$E_{rest}(M, m) = Mc^2 + mc^2 - \frac{1}{2}\sqrt{mMc^2}$$

The whole energy such a “binary” system is as follows:

$$\begin{aligned} E(M, m) &= Mc^2 + mc^2 - \frac{1}{2}\sqrt{mM}c^2 \\ &= (M + m)c^2 \left\{ 1 - \frac{1}{2}\sqrt{\frac{\mu}{M + m}} \right\}, \end{aligned}$$

where $\mu = mM/(m + M)$.

$$Mc^2, \quad mc^2, \quad (M + m)c^2 - \text{rest energy}$$

$$\Delta E = \frac{1}{2}\sqrt{\mu(M + m)}c^2 = \frac{1}{2}\sqrt{mM}c^2 - \text{energy corresponding}$$

$$\text{the “mass defect” } \Delta m = \frac{1}{2}\sqrt{mM}$$

“Gravitational fusion” picture

$$\begin{aligned}\lambda &= \frac{[\text{eliminated energy}]}{[\text{rest energy}]} = \frac{\Delta E}{(M + m)c^2} \\ &= \frac{\sqrt{mM}}{2(M + m)} = \frac{1}{2} \frac{\sqrt{\delta}}{1 + \delta},\end{aligned}$$

where

$$\delta = \frac{m}{M}, \quad \delta_0 \leq \delta \leq 1, \quad \delta_0 = \frac{M_{min}}{M_{max}}$$

$$\text{Maximum } \lambda_{max} = \frac{\Delta E_{eliminated}}{E_{rest}} = \frac{1}{4}$$

$$\text{under } \delta = 1 \quad (m = M)$$

Such a “fusion” of two equal rest mass M lead to mass $= 3M/2$, the Energy $\Delta E = Mc^2/2$ eliminate (**3 : 1**).

Extended Phase Space as a Basic Manifold

Space of states (QTP) and Energy (H) (sec [J. Synge, 1962]) in general: $2 + 2N$ dimensions. The metric is

$$\eta_{AB} = \text{diag}\{\overbrace{1, -1, -1, \dots, -1}^N\}$$

(X, Y) – pair of conjugated pseudoeuclidean vectors $X_A, Y_A, A, B = 0, 1, \dots, N$ ($N + 1$ dimension).

The Poisson brackets for $\varphi(X, Y), f(X, Y)$,

$$\{\varphi, f\}_{A,B} = \sum_{k=0}^N \left\{ \frac{\partial \varphi}{\partial X^k} \frac{\partial f}{\partial Y_k} - \frac{\partial \varphi}{\partial Y^k} \frac{\partial f}{\partial X_k} \right\}$$

so that

$$\{X_A, Y_A\}_{A,B} = \eta_{AB}$$

Dynamic System in (QT, PH) is defined, when the [Energy Surface](#) ($2N + 1$ dimation)

$$\Omega(X, Y) = 0$$

is defined. Solving this equation on Y_0 we obtain [Energy equation](#)

$$Y_0 = F(X_k, Y_l, X_0) \quad k, l = 1, 2, \dots, N$$

The corresponding Hamilton function of the System is [by def:](#)

$$\mathcal{H} \equiv Y_0 = \mathcal{H}(X_k, Y_l, Y_0)$$

We shall consider $(2 + 2 \cdot 3)$ – dimensional version (QT, PH)

$$A, B \rightarrow \mu, \nu = 0, 1, 2, 3, \quad \eta_{\mu\nu} = \text{diag}\{1, -1, -1, -1\},$$

X_μ, X_μ are real 4-dimensional pseudoeuclidean $SO(3, 1)$ – vectors.

Complexification of (QT, PH)

Let us introduce the complex 4-vector Z^μ according definition:

$$Z^\mu = X^\mu + iY^\mu, \quad \eta_{\mu\nu} = \text{diag}\{1, -1, -1, -1\}$$

and define the real norm

$$\begin{aligned} Z^\mu Z_\mu^* &= Z^\mu \eta_{\mu\nu} Z^{\nu*} \\ &= |Z_0|^2 - |Z_1|^2 - |Z_2|^2 - |Z_3|^2 = \text{inv} \end{aligned}$$

Complex Lorentz Group with a real metric (Barut Group [Barut 1964])

The group of 4×4 complex matrices Λ of transformations

$$Z' = \Lambda Z, \quad Z = X + iY$$

satisfying the condition

$$\Lambda \eta \Lambda^\dagger = \eta, \quad (\dagger - \text{hermitien conjugation})$$

Notice: the diagonal matrices Λ_R and $\bar{\Lambda}_R \in (BG)$

$$\Lambda_R = -iI_4 = -i\text{diag}\{1, 1, 1, 1\}$$

and

$$\bar{\Lambda}_R = -i\eta_{\mu\nu} = -i\text{diag}\{1, -1, -1, -1\}$$

Metric invariant (no positive-defined)

$$Z^\mu Z_\mu^* = |Z_0|^2 - |Z_1|^2 - |Z_2|^2 - |Z_3|^2 = \text{inv}$$

Three possibilities: $|Z|^2 \geq 0$, $|Z|^2 = 0$ – isotropic case

$SU(3,1)$ – subgroup structure

(QTPH)

$$QT = \{\underline{r}, ct\} = x^\mu$$

$$PH = \{\underline{p}, E/c\} = p^\mu$$

$$\vec{\beta}_v = \vec{v}/c$$

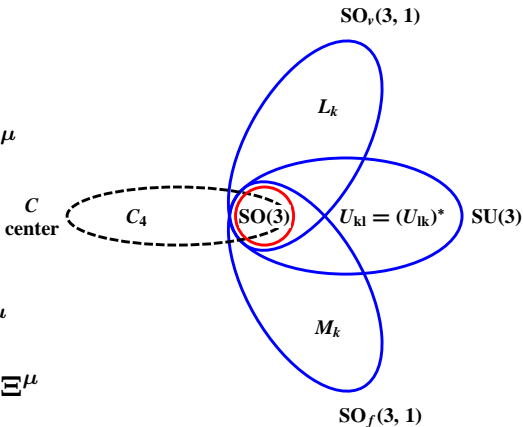
(PTQH)

$$PT = \{\underline{p}, F_0 t\} = \Pi^\mu$$

$$QH = \{\underline{r}, E/F_0\} = \Xi^\mu$$

$$\vec{\beta}_f = \vec{f}/F_0$$

$$F_0 = F_{max}$$



$$k, l = 1, 2, 3$$

Reciprocal Invariant S_B^2 as BG metric invariant

$Z^\mu Z_\mu^*$

By def: $Z_\mu = X_\mu + iY_\mu$

$$X_\mu = \frac{1}{q_e} x_\mu = \frac{1}{q_e} \{x_0, x_k\}, \quad Y_\mu = \frac{1}{p_e} p_\mu = \frac{1}{p_e} \{p_0, p_k\}.$$

Then:

$$\begin{aligned} S_B^2 &= \frac{1}{q_e^2} \left(x_0^2 - \sum_{k=1}^3 x_k^2 \right) + \frac{1}{p_e^2} \left(p_0^2 - \sum_{k=1}^3 p_k^2 \right) = \\ &= \frac{x_0^2}{q_e^2} + \frac{p_0^2}{p_e^2} - \sum_{k=1}^3 \left(\frac{x_k^2}{q_e^2} + \frac{p_k^2}{p_e^2} \right) = |Z_0|^2 - \sum_{k=1}^3 |Z_k|^2 = Z^\mu Z_\mu^* \end{aligned}$$

Let x_μ, p_μ to be 4-vectors under $SO_v(3, 1)$ -transformations

Notice! There exist alternative:

$$\bar{Z}_\mu = \bar{X}_\mu + i\bar{Y}_\mu$$

where

$$\bar{X}_\mu = \Xi_\mu \stackrel{\text{def}}{=} \left\{ \frac{x_0}{q_e}, \frac{p_k}{p_e} \right\}, \quad \bar{Y}_\mu = \Pi_\mu \stackrel{\text{def}}{=} \left\{ \frac{p_0}{p_e}, \frac{x_k}{q_e} \right\}.$$

It is evident:

$$\begin{aligned} |\bar{Z}_\mu|^2 &= \Xi^\mu \Xi_\mu + \Pi^\mu \Pi_\mu = \frac{x_0^2}{q_e^2} - \sum_{k=1}^3 \frac{p_k^2}{p_e^2} + \frac{p_0^2}{p_e^2} - \sum_{k=1}^3 \frac{x_k^2}{q_e^2} = \\ &= \frac{1}{q_e^2} x^\mu x_\mu + \frac{1}{p_e^2} p^\mu p_\mu = |Z_\mu|^2. \end{aligned}$$

The natural supposition: Ξ_μ and Π_μ are 4-vectors under $SO_f(\mathbf{3}, \mathbf{1})$ -transformations

Selfreciprocal Case: $S_B^2 = 0$

Quadratic form $S_B^2 = Z^\mu Z_\mu^*$ is not positive defined

Three possibilities $Z^\mu Z_\mu^* = 0$, $Z^\mu Z_\mu^* \gtrless 0$.

Especially interesting is:

$$S_B^2 = \frac{1}{p_e^2} p^\mu p_\mu + \frac{1}{q_e^2} x^\mu x_\mu = 0$$

Two possibilities

①

$$p^\mu p_\mu = x^\mu x_\mu = 0$$

(light cone, massless particles, standard clock synchronisation)

②

$$x^\mu x_\mu = -q_e^2, \quad p^\mu p_\mu = p_e^2,$$

$$q_e = r_g(m_e) = \frac{2m_e G_N}{c^2}, \quad p_e = m_e c.$$

BG: some kinematic outcome

(1) Reciprocal-Invariant time T_R

$$z_\mu = x_\mu + i\mathfrak{x}_0^{-1}p_\mu, \quad \mathfrak{x}_0 = \frac{p_e}{q_e} = \frac{c^3}{4G_N}$$

$$\begin{aligned} dT_R &\stackrel{\text{def}}{=} c^{-1} \sqrt{z^\mu z_\mu^*} \\ &= c^{-1} \sqrt{dx_0^2 - d\underline{r}^2 + \mathfrak{x}_0^{-2}(dp_0^2 - d\underline{p}^2)} \\ &= dt \sqrt{1 - \underline{\beta}_v^2 - \underline{\beta}_f^2 + \underline{\beta}_v^2 \underline{\beta}_f^2 \cos^2 \alpha} \\ &= dt \sqrt{\left(1 - \underline{\beta}_v^2\right) \left(1 - \underline{\beta}_f^2\right) - \left(\underline{\beta}_f \times \underline{\beta}_v\right)^2}. \end{aligned}$$

where

$$\underline{\beta}_v = \frac{\dot{\underline{r}}}{c}, \quad \underline{\beta}_f = \frac{\dot{\underline{p}}}{F_G}, \quad F_G = \frac{c^4}{4G_N}, \quad (\cdot) = \frac{d}{dt},$$

t – is the laboratory time (comoving observer time).

α – is the angle between $\dot{\underline{r}}$ and $\dot{\underline{p}}$.

(2) In the comoving \mathbf{RF} :

$$dT_R(\beta_v = 0) = dt\sqrt{1 - \underline{\beta}_v^2}$$

There must exist an extra time retardation.

$$(3) \cos \alpha = \pm 1, \quad (\underline{\beta}_f \times \underline{\beta}_v) = 0$$

$$dT_R^{\parallel} = dt\sqrt{1 - \underline{\beta}_v^2}\sqrt{1 - \underline{\beta}_f^2}$$

$$(4) \cos \alpha = 0, \quad (\underline{\beta}_f \times \underline{\beta}_v)^2 = \underline{\beta}_f^2 \cdot \underline{\beta}_v^2$$

$$dT_R^{\perp} = dt\sqrt{1 - \underline{\beta}_v^2 - \underline{\beta}_f^2} \neq dT_R^{\parallel}$$

$$(5) \cos \alpha = \pm\sqrt{2}/2,$$

$$dT_R^{\parallel} = dT_R^{\perp}$$

(6) The upper limit of $|\dot{\underline{r}}|$ and $|\dot{\underline{p}}|$ in the expression for dT_R :

$$|\dot{\underline{r}}|_{max} = c/\sqrt{2}, \quad |\dot{\underline{p}}|_{max} = F_G/\sqrt{2}$$

Complex Lorentz Boost Transformations

Let us represent a **BG**-transformations using the following parametrization:

$$\Lambda(\beta) = \begin{pmatrix} \Gamma & \underline{\Gamma} \cdot \underline{\beta} \\ \underline{\Gamma} \cdot \underline{\beta}^* & I_3 + \frac{\underline{\beta} \cdot \underline{\beta}^*}{|\underline{\beta}|^2} (\Gamma - 1) \end{pmatrix}, \quad \Lambda(\beta) \eta \Lambda^\dagger(\beta) = \eta$$

where $\underline{\beta} = \underline{\beta}_v + i\underline{\beta}_F = \underline{v}/c + i\underline{F}/F_G$,

\underline{v} – is a const velocity

\underline{F} – is a const force, $F_G = c^4/4G_N$ – max force

$\Gamma = \det \Lambda^{-1}(\beta)$, $I_3 = \text{diag}\{1, 1, 1\}$, $\underline{\beta} \cdot \underline{\beta}^* = 3 \times 3$ tensor-diada.

$$Z' = \Lambda(\beta) Z : \quad \begin{pmatrix} Q'_0 + iP'_0 \\ \underline{Q}' + i\underline{P}' \end{pmatrix} = \Lambda(\beta) \begin{pmatrix} Q_0 + iP_0 \\ \underline{Q} + i\underline{P} \end{pmatrix}$$

The simplest case $\underline{\beta}_v = \{\beta_v, 0, 0\}$, $\underline{\beta}_F = \{\beta_F, 0, 0\}$

$$\Lambda(\beta) = \Gamma(\beta) \begin{pmatrix} 1 & \beta \\ \beta^* & 1 \end{pmatrix}, \quad \beta = \beta_v + i\beta_F,$$

$$\Gamma(\beta) = \frac{1}{\sqrt{1 - |\underline{\beta}|^2}}, \quad \beta_v = \frac{v}{c}, \quad \beta_F = \frac{F}{F_G},$$

F – is the constant force uniformly accelerated (noninertial)
 FR .

$$\begin{pmatrix} Q'_0 + iP'_0 \\ Q' + iP' \end{pmatrix} = \Gamma(\beta) \left\{ \begin{pmatrix} 1 & \beta_v \\ \beta_v^* & 1 \end{pmatrix} + i \begin{pmatrix} 0 & \beta_F \\ -\beta_F & 0 \end{pmatrix} \right\} \begin{pmatrix} Q_0 + iP_0 \\ Q + iP \end{pmatrix}$$

$$\Gamma(\beta) = (1 - \beta_v^2 - \beta_F^2)^{-1/2}$$

$$\beta_F = 0$$

$$= \gamma_v \begin{pmatrix} ct' \\ x' \end{pmatrix} = \gamma_v \begin{pmatrix} 1 & \beta_v \\ \beta_v & 1 \end{pmatrix} \begin{pmatrix} ct \\ x \end{pmatrix},$$

$$= \gamma_v \begin{pmatrix} E'/c \\ p' \end{pmatrix} = \gamma_v \begin{pmatrix} 1 & \beta_v \\ \beta_v & 1 \end{pmatrix} \begin{pmatrix} E/c \\ p \end{pmatrix},$$

$$\gamma_v = 1/\sqrt{1 - \beta_v^2}$$

$$\dot{x}' = \frac{\dot{x} + V}{1 + \dot{x}V/c^2} :$$

$$\beta'_v = \frac{\beta_v + \beta_V}{1 + \beta_v\beta_V}$$

$$\beta_v = 0, \quad F_G = \frac{c^4}{4G_N}$$

$$= \gamma_F \begin{pmatrix} t' \\ p'/F_G \end{pmatrix} = \gamma_F \begin{pmatrix} 1 & -\beta_F \\ -\beta_F & 1 \end{pmatrix} \begin{pmatrix} t \\ p/F_G \end{pmatrix},$$

$$= \gamma_F \begin{pmatrix} E' \\ F_G x' \end{pmatrix} = \gamma_F \begin{pmatrix} 1 & -\beta_F \\ -\beta_F & 1 \end{pmatrix} \begin{pmatrix} E \\ F_G x \end{pmatrix},$$

$$\gamma_F = 1/\sqrt{1 - \beta_F^2}$$

$$\dot{p}' = \frac{\dot{p} - F}{1 - \dot{p}F/F_G^2} :$$

$$\beta'_f = \frac{\beta_f - \beta_F}{1 - \beta_f\beta_F},$$

$$\dot{p} = \frac{dp}{dt} = f, \quad \frac{dE}{dx} = -f$$

Hamiltonian selfreciprocal one-particle classical model

The selfreciprocal expression of Born interval

$$S_B^2 = (x_0^2 - \underline{r}^2) / q_e^2 + (p_0^2 - \underline{p}^2) = 0$$

is **by definition the Energy equation** in the
(**2 + 2 · 3**)-dimensional extended Phase Space (QTPH)
[J.Synge, Class. Dyn.] and

$$\begin{aligned} & P_0^2 + T^2 - (\underline{P}^2 + \underline{Q}^2) \\ &= \left(P_0 + \sqrt{\underline{P}^2 + \underline{Q}^2} \right) \left(P_0 - \sqrt{\underline{P}^2 + \underline{Q}^2} \right) = 0 \end{aligned}$$

is the **energy surface equation**. According to (**QTPH**)
definition $\mathcal{H}(\underline{Q}, \underline{P}, T) \equiv -P_0$ is the Hamilton function of
system. In classical version we are to choose the positive sign.

So that Hamilton function $\mathcal{H}(Q_k, P_k, T)$ is as follows

$$\mathcal{H}(Q_k, P_k, T) = \sqrt{H_0^2 - T^2}, \quad Q_k = \frac{x_k}{q_e}, \quad P_k = \frac{p_k}{p_e},$$

where

$$T = x_0/q_e = v_e t \quad (v_e = v(m_e) = \mathfrak{x}_0/m_e)$$

$$H_0 = \sqrt{P_k^2 + Q_k^2} - \text{energy (integral of motion)}$$

$$\frac{dH_0}{dT} = \{H_0(\underline{P}, \underline{Q}), \mathcal{H}(\underline{P}, \underline{Q}, T)\}_{q,p} = 0$$

The Hamiltonian equations

$$\frac{dQ_k}{dT} = \frac{P_k}{\sqrt{H_0^2 - T^2}}, \quad \frac{dP_k}{dT} = \frac{-Q_k}{\sqrt{H_0^2 - T^2}}$$

possess under the spherical-symmetric initial conditions the following solution

$$|\vec{Q}| = \sin(\varphi + \varphi_{min}), \quad |\vec{P}| = \cos(\varphi + \varphi_{min}),$$

$$\varphi = \arcsin(T/H_0) \quad (0 < \varphi < \pi/2),$$

$$\varphi_{min} \neq 0 - \text{minimal angle} \sim \text{minimal value of } p_e q_e$$

The model: massive pulsating sphere

Frequency of pulsation (comoving FR time)

$$\omega_0(m_e) = \frac{\mathfrak{a}_0}{H_0 m_e} = \frac{c^3}{4m_e G_N} \sim \frac{10^{38} \text{s}^{-1}}{m_e}.$$

Pulsation Period

$$T_0(m_e) = 2\pi/\omega_0(m_e) \sim m_e \cdot 10^{-38} \text{s}.$$

Spatial peak value

$$R(m_e) = 4m_e G_N / c^2 \sim m_e \cdot 10^{-28} \text{sm}.$$

For $m_e = m_u \sim 10^{55} \text{ gr}$,

$$T(m_u) \sim 10^{17} \text{s} \sim H_0^{-1} \sim \Delta t_u, \quad R(m_u) \sim 10^{27} \text{sm} \sim R_u$$

.

Elliptic trajectory: $r - p$ -space

The eccentricity

$$\varepsilon(\alpha) = \frac{\sqrt{\vec{Q}_+^2 - \vec{Q}_-^2}}{|\vec{Q}_+|} = \frac{\sqrt{2} (1 - \alpha^2)^{1/4}}{(1 + \sqrt{1 - \alpha^2})^{1/2}}$$

$$\varepsilon(\alpha)|_{\alpha \ll 1} \cong 1 - \frac{\alpha^2}{8} = 1 - \frac{\vec{Q}^2 \vec{P}^2}{2(\vec{Q}^2 + \vec{P}^2)^2}, \quad \text{but } \neq 1.$$

$$\alpha^2|_{P \ll Q} \cong \frac{4\vec{P}^2}{\vec{Q}^2} = \frac{2p}{r\mathfrak{a}_0}, \quad \frac{p}{r} \ll \mathfrak{a}_0$$

$P(\dot{Q})$ – dependence

$$\dot{Q} = \frac{d\vec{Q}}{dT} = \frac{\vec{P}}{\sqrt{\vec{P}^2 + \vec{Q}^2 - T^2}} \rightarrow \vec{P}^2 = \frac{\dot{Q}^2}{1 - \dot{Q}^2} (\vec{Q}^2 - T^2)$$

$$\vec{Q}^2 - T^2 = 1 - \text{hyperbolic motion!}$$

$$\vec{P}_{max}^2 = \frac{\dot{Q}_{max}^2}{1 - \dot{Q}_{max}^2}. \quad \text{But} \quad |\vec{P}_{max}| = 1$$

Therefore $2\dot{Q}_{max}^2 = 1$ and $\dot{r}_{max} = c/\sqrt{2}$

The case $S_B^2 = \Lambda_B \neq 0$

From $S_B^2 = P^\mu P_\mu + Q^\mu Q_\mu = \Lambda_B > 0$ to
 $\bar{S}_B^2 = S_B^2 - \Lambda_B = 0$.

The Energy Integral is:

$$H_0(P, Q)|_{\Lambda_B=0} \rightarrow H_0(P, Q, \Lambda_B)$$

$$= H_0 = \sqrt{\vec{P}^2 + \vec{Q}^2 + \Lambda_B}$$

$$p_e(m), \quad q_e(m), \quad \frac{p_e(m)}{q_e(m)} = \mathfrak{x}_0$$

$$\rightarrow p_e(M), \quad q_e(M), \quad \frac{p_e(M)}{q_e(M)} = \mathfrak{x}_0$$

Scale transformation of the parameter: $M = m\sqrt{1 + \Lambda_B}$.

Consider the case $\Lambda_B = 1$.

Two versions: Classical and Quantum

The Newtonian Limit

Let us consider the Energy integral $H_0 = \sqrt{\underline{P}^2 + \underline{Q}^2 + 1}$
under conditions $|\underline{Q}| \ll 1$, $|\underline{P}| \ll 1$, $|\underline{P}|/|\underline{Q}| \ll 1$.
Use the dimensional values

$$\begin{aligned} E &= cp_e H_0 = cp_e \left(\frac{p^2}{p_e^2} + \frac{r^2}{q_e^2} + 1 \right)^{1/2} \\ &= cp_e \sqrt{1 + \frac{r^2}{q_e^2}} \left\{ 1 + \frac{p^2}{p_e^2} \left(1 + \frac{r^2}{q_e^2} \right)^{-1} \right\}^{1/2} \\ &\approx mc^2 + \frac{1}{2}mv^2 + \frac{1}{2}mr^2\omega_0^2(m). \end{aligned}$$

Omitting the “rest energy” mc^2 we receive

$$E_N = \frac{1}{2}mv^2 + \frac{1}{2}I_r\omega_0^2(m)$$

where $I_r = mr^2$ – moment of inertia

$\omega_0 = \mathfrak{a}_0/m$ – angular velocity

$E_{rot} = mr^2\omega_0^2/2$ – is the kinesthetic energy of rotation.

Canonical Quantization $S_B^2 \rightarrow \hat{S}_B^2$

According to the universal condition

$$m_e^2 = \frac{ca_e}{4\pi G_N} \rightarrow \left(a_e = \frac{h}{2}\right) \rightarrow \frac{\hbar c}{4G_N} = \frac{1}{4}M_P^2$$

the parameters \mathbf{p}_e , \mathbf{q}_e , \mathbf{E}_e , \mathbf{t}_e becomes Planck. The basic Born equation can be written in the form (Fock presentation)

$$\hat{S}_B |n_0, n\rangle = 2 \left(\hat{H}_0 - \sum_{k=1}^3 \hat{H}_k \right) |n_0, n\rangle = \lambda_B(n_0, n) |n_0, n\rangle,$$

where \hat{H}_μ ($\mu = 0, 1, 2, 3$) are linear *QHOSc* Hamiltonians.

The discrete spectrum of (mass)²:

$$M^2(n_0, n) = \{n_0 - (n + 1)\} M_P^2 = N(n_0, n) \left(\frac{\hbar c}{G_N} \right)$$

where $N(n_0, n) = n_0 - (n + 1)$, $n = \sum_{k=1}^3 n_k$.

Each of $n_\mu = 0, 1, 2, \dots$ ($\mu = 0, 1, 2, 3$) independently.

Eigenstates

$$|n_0, n\rangle = \prod_{\mu=0}^3 |n_\mu\rangle, \quad N(n) = \frac{1}{2}(n+1)(n+2)$$

is the degeneracy degree of $|n\rangle \sim |n_1\rangle|n_2\rangle|n_3\rangle$ states. In reality the operator \hat{S}_B^2 is looking out as the quantum-mechanical action operator having linear spectrum S_N in unit of \hbar :

$$S_N = N\hbar \quad (N = 1, 2, 3, \dots)$$

Canonical Quantization

$$H_0(\underline{P}, \underline{Q}) \rightarrow \hat{H}_0(\hat{\underline{P}}, \hat{\underline{Q}})$$

Let us consider the variables \hat{P}_k, \hat{Q}_l in the classical Energy integral

$$H_0 = \sqrt{\underline{P}^2 + \underline{Q}^2 + 1}$$

as canonically conjugated operators $[\hat{P}_k, \hat{Q}_l] = -i\delta_{kl}$, the constants p_e and q_e in definition

$$\hat{Q}_l = \frac{\hat{x}_l}{q_e}, \quad \hat{P}_k = \frac{\hat{p}_k}{p_e}$$

are Planck's.

$$\hat{H}_0 \rightarrow \hat{H}_0 = \left(\hat{\underline{P}}^2 + \hat{\underline{Q}}^2 + 1 \right)^{1/2}$$

The root extraction according Dirac procedure gives (in noncovariant notations):

for **DO** Hamiltonian and its Supersymmetry Partner H_{DO}^{\pm}

$$\begin{aligned}\hat{H}_0^{\pm}(\hat{P}, \hat{Q}, 1) &= \alpha_k \hat{P}_k + \beta \left(I \mp i \alpha_k \hat{Q}_k \right) = \\ &= \hat{H}_{DO}^{\pm} = \alpha_k \left(\hat{P}_k \mp i \beta \hat{Q}_k \right) + \beta,\end{aligned}$$

where α_k, ρ_k ($k = 1, 2, 3$) – 4×4 Dirac matrices, $\beta = \rho_3$.

$\hat{H}_{DO}^{\pm} = \hat{H}_{DO}^{\pm}(M_{Pl})$ – well known Dirac Oscillator

Hamiltonian which describes the one-half Spin particle with a Plank mass. The exact solutions of this model are well known. But generally accepted physical interpretation of **DO**-model is absent up to now.

Thank you for your attention